Beam dynamics limits for Low-Energy RHIC ("critRHIC" project)

A. Fedotov, I. Ben-Zvi, X. Chang, D. Kayran, V. Litvinenko, E. Pozdeyev, T. Satogata

HB2008 Workshop, Nashville, Tennessee, August 25-29, 2008



Low-energy RHIC operation

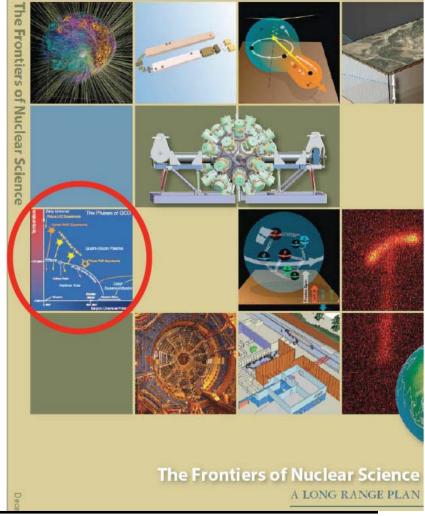
There is substantial and growing interest in RHIC heavy ion collisions with c. m. energy in the range $\sqrt{s_{NN}}$ = 5-50 GeV/n

RIKEN workshop (BNL, March 2006):

"Can we discover the QCD critical point at RHIC?"

Suggested energy scan: $\sqrt{s_{NN}} = 5$, 6.3, 7.6, 8.8, 12.3, 18, 28 GeV/n

Three test runs were done in 2006, 2007, 2008 at low-energies in RHIC (Todd Satogata et al., PAC07; RHIC 2008 retreat).

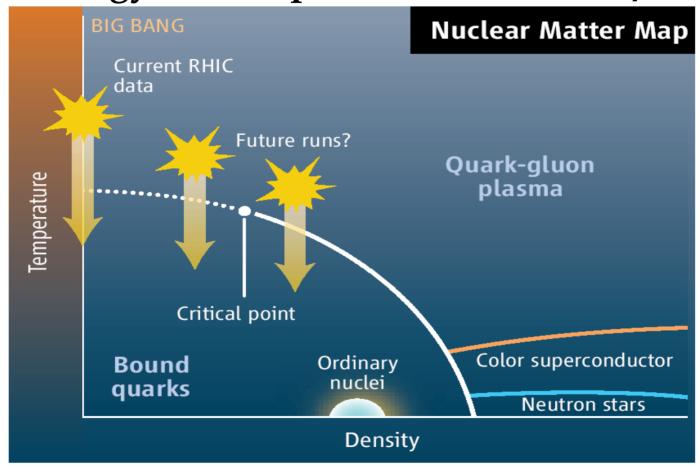






A. Cho, Science, 312 (14 Apr 2006)

Low-energy RHIC operation: 2.5-25 GeV/n

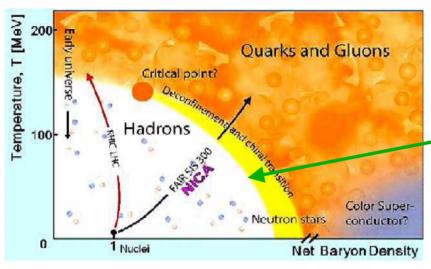


Landmark study. Physicists have seen a smooth transition from bound quarks to quark-gluon plasma (dotted line). They now hope to find the point beyond which the transition becomes violent (white line).



A. Fedotov et al.





High statistics/ luminosity is needed.

deconfinement and/or chiral symmetry restoration phase transitions

the first stage

- Multiplicity and global characteristics of identified hadrons including multi-strange particles
- Fluctuations in multiplicity and transverse momenta
- Directed and elliptic flows for various hadrons
- HBT and particle correlations

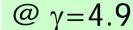
the second stage

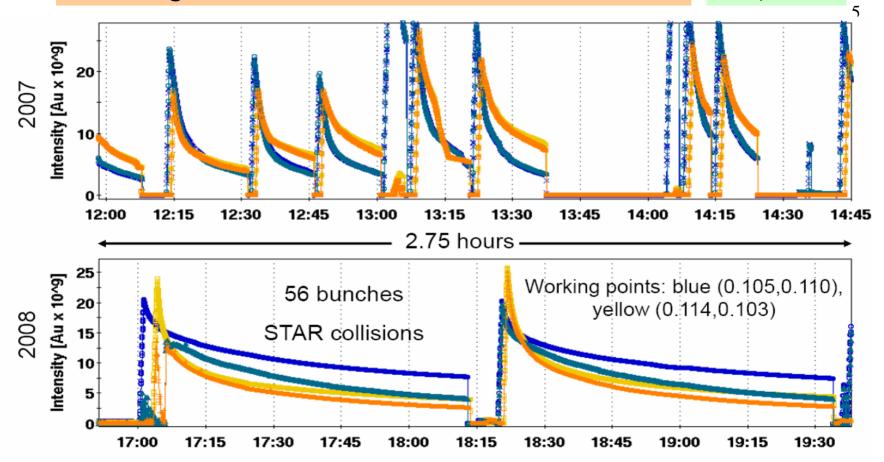
Electromagnetic probes (photons and dileptons)





T. Satogata (RHIC retreat, March 31, 2008)





- 2008 blue beam lifetime: 3.5 minutes (fast), 50 minutes (slow)
- Sextupole reversal and elimination of octupoles clearly helped beam lifetime
- Injection efficiency and yellow beam lifetime can clearly benefit from further tuning



Summary of 2008 low-energy test run

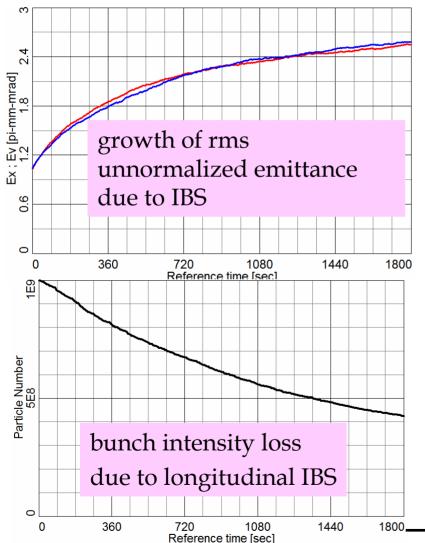
■ $\sqrt{s_{NN}}$ = 9 GeV

- T. Satogata (RHIC retreat, March 31 2008)
- All setup worked very well with h=366; commendations to setup team!
- Defocusing sextupole reversal, octupole removal improved blue lifetime by x4!
- ~50-60% injection efficiency
- STAR collisions about 13h after first beam; PHENIX about 24h after first beam.
- Clean vernier scans, unambiguous collisions achieved in both experiments
- Experiment useful event rates: 0.7-1Hz with 56 bunches, 0.4-0.5e9/bunch
 - Luminosity max 3.5x10²³ cm⁻² s⁻¹, average 1.2x10²³ cm⁻² s⁻¹
- Problems: cogging, AGS-RHIC synchro, kicker timing, PHENIX collision signal, intensity dynamic range for IBS measurements
- $\sqrt{s_{NN}}$ = 5 GeV
 - Timing setup worked very well with h=387; more commendations!
 - 10% injection efficiency, very nonlinear lattice (main dipole b₂ large)
 - Problems: b4-dh0 ps failure, limited bunched beam, nonworking orb correction due to linear model





IBS for Au ions in RHIC for lowest energy point7



Simulation parameters

| Parameters | Value | | |
|--|--------|--|--|
| Kinetic energy of Au ions, GeV/nucleon | 1.57 | | |
| Relativistic γ | 2.68 | | |
| Bunch intensity, 109 | 1.0 | | |
| Rms momentum spread | 4×10-4 | | |
| Rms bunch length, cm | 155 | | |
| Rms emittance (unnormalized), µm | 1.04 | | |
| RF harmonic | 387 | | |
| RF voltage, kV | 300 | | |

BROOKHAVEN

A. Fedotov et al.



Beam dynamics luminosity limits

IBS:

• Strong IBS growth - can be easily counteracted by Electron Cooling system.

Beam-beam:

• Becomes significant limitation for RHIC parameters only at $\gamma > 10$.

Space-charge::

• At lowest energies, strongest limitation on achievable ion beam peak current is expected to be given by space-charge effects. This prohibits application of strong electron cooling.





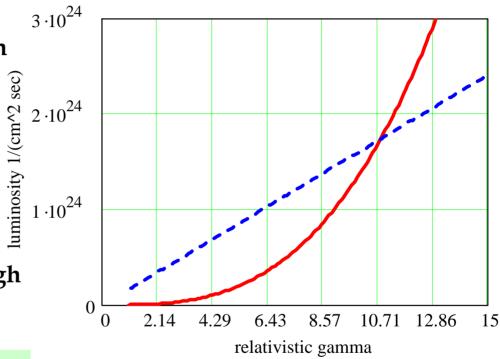
Luminosity limitation by space-charge and beam-beam 9

Luminosity expressed through beam-beam parameter ξ:

$$L = \frac{A}{Z^2 r_p} \frac{N_i c}{\beta^* C} \frac{2\gamma \beta^2}{1 + \beta^2} f\left(\frac{\sigma_s}{\beta^*}\right) \xi$$

Luminosity expressed through space-charge tune shift ΔQ :

$$L = \frac{A}{Z^{2} r_{p}} \frac{N_{i} c}{\beta^{*}} \frac{\sqrt{2\pi} \sigma_{s}}{C^{2}} \gamma^{3} \beta^{2} f\left(\frac{\sigma_{s}}{\beta^{*}}\right) \Delta Q$$



Blue dash line: beam-beam limitation with beam-beam parameter 0.005 per IP.

Red: space-charge limitation with $\Delta Q = 0.05$

What is acceptable space-charge tune shift for long life time in RHIC with collisions?

This question could be explored by measurement during the test runs at low-energy.

At injection energy, with space-charge tune shift of ΔQ =0.01, life time in RHIC is good.

In June 2007 test run with ΔQ =0.03 the life time was bad although it seems to be related to other effects.

- for example, in SPS, life time of few minutes with ΔQ =0.1 was reported
- in LEAR, electron cooling allowed to increase ΔQ to 0.1 in operation
- Presently, we assume that ΔQ =0.05-0.07 as a limit for RHIC (since we would like to have long life time in RHIC minutes). Perhaps, cooling can help to operate at slightly larger space-charge tune shift.





Incoherent space-charge tune shift for lowest energy points in RHIC

| γ | h | ε _{95%,n} | N, ×10 ⁹ | $\Delta Q_{sc,G}$ | $\Delta Q_{\mathrm{sc,fb}}$ |
|------|----------------------|--------------------|-----------------------|-------------------------|---------------------------------------|
| | (rf harmonic number) | μm | (ion bunch intensity) | (Gaussian distribution) | (full bucket, parabolic distribution) |
| 2.67 | 387 | 15 | 0.5 | 0.12 | 0.07 |
| 3.37 | 375 | 15 | 1 | 0.15 | 0.08 |
| 4.41 | 369 | 15 | 1 | 0.1 | 0.05 |
| 4.7 | 366 | 15 | 1 | 0.07 | 0.04 |
| 6.6 | 360 | 15 | 1 | 0.04 | 0.02 |





Possible scenarios under constraint of space-charge limit₂

- 1. Colliding coasting beam would reduce space-charge tune shift and, with electron cooling, would allow significant luminosity increase. However, present detector system is not set for such approach. Discussion with physicists is in progress.
- 2. (working scenario)

Operating with bunched beam at space-charge limit with 110 bunches in the machine - favored by the experiments, since they require 100 nsec bunch spacing for now.



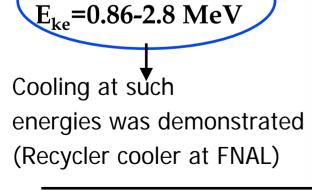
RHIC low-energy electron cooler energies of interest

The lowest energy points from proposed energy scan ion kinetic energies: E_{ki} =1.57, 2.2, 2.9, 3.45, 5.2 GeV/nucleon can benefit the most from electron cooling.

These corresponds to electron beam

2007 estimates by T. Satogata

kinetic energies:



| √s _{nn} [GeV] | μ _B [MeV] | <bbc rate=""> [Hz]</bbc> | Days/ Mevent | # events | # beam days |
|---------------------------|-------------------------|------------------------------|-----------------|----------|------------------|
| 4.6 | 570 | 3 (~5) | 9 (3) | 5M | 45 (15+ N |
| 6.3 | 470 | 7 (~50) | 4 (0.3) | 5M | 20 (3+1) |
| 7.6 | 410 | 13 (~150) | 2 (0.1) | 5M | 10 (1+1) |
| 8.8 | 380 | 20 (300) | 1.5 (<1) | 5M | 7.5 (1+1) |
| 12 | 300 | 54 (~1000) | 0.5 (<1) | 5M (>5M) | 2.5 (1+1) |
| 18 | 220 | >100 (>1000) | 0.25 (<1) | 5M (>5M) | 1.5 (1+1) |
| 28 | 150 | >100 (>1000) | 0.25 (<1) | 5M (>5M) | 1.5 (1+2) |





Options for the cooler

1. DC cooler:

- 1.1 Using Recycler (FNAL) e-cooler when it becomes available.
- 1.2 Building new 3MeV DC electron cooler.

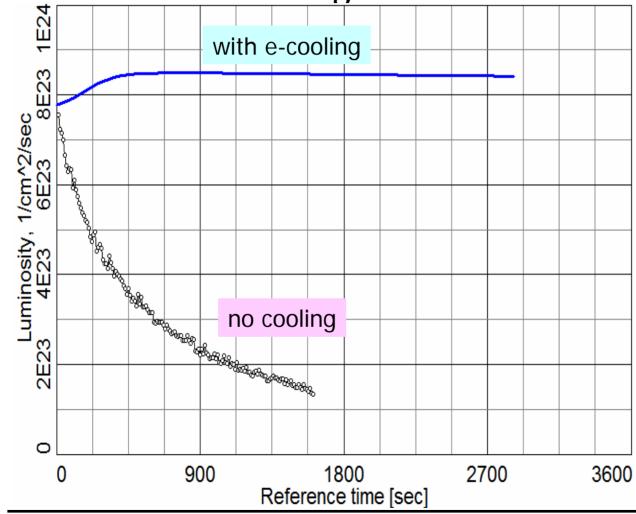
2. RF gun cooler:

- 2.1 Using 703 MHz gun and cooling with a pulse of electron bunches (20 low-charge 50pC bunches per pulse) to reduce space-charge effects is a single bunch (studies by D. Kayran et al.).
- 2.2 Using 56 MHz cavity to produce long electron bunch with required parameters for cooling (studies by X. Chang et al.).





γ =2.7 – lowest energy point. Luminosity with and without electron cooling.

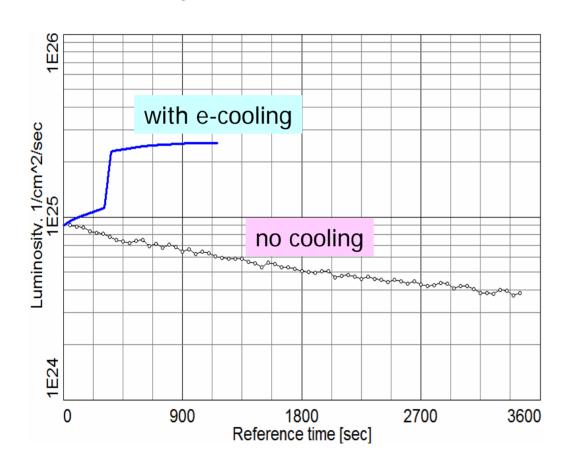




A. Fedotov et al.

γ =6.6 - highest energy point of the cooler planned.

Luminosity with and without electron cooling.



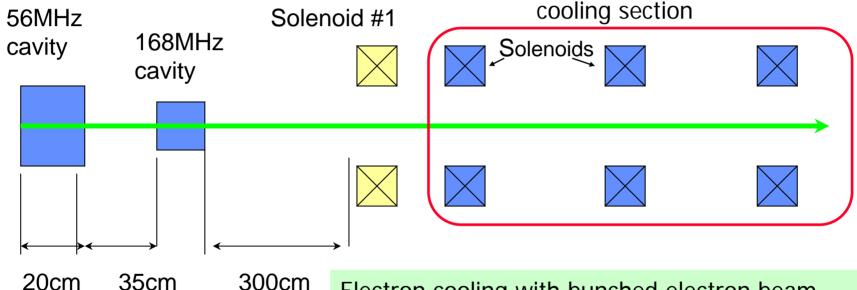
At higher energies, by providing sufficient cooling we can cool emittance of ion beam until space-charge limit, which in turn allows to decrease β^* .

In example shown, electron cooling provides a factor of 6 improvement in integrated luminosity.





RF gun cooler based on 56MHz gun (X. Chang et al.)



Electron cooling with bunched electron beam produced by 56MHz cavity – long bunch is needed to minimize space-charge effect in beam transport at low-energies.





Result of electron beam simulations using 56 MHz gun

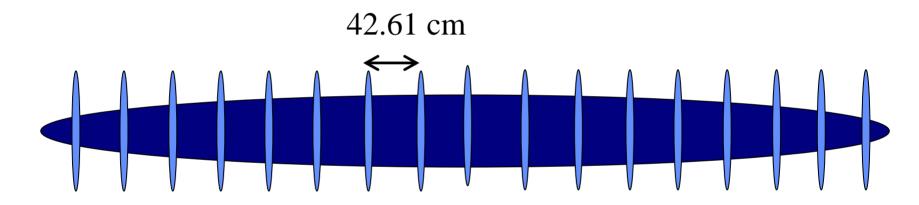
| Kinetic Energy (MeV) | Charge (nC) | Rms parameters | Requirement | Simulation |
|----------------------------|----------------|---------------------|--------------------|---|
| | | | | |
| | | ΔΕ/Ε | 5×10 ⁻⁴ | ~5×10 ⁻⁴ @80% core (0.8nC) |
| | 1 | emittance (mm.mrad) | ~ 3 | ~3.2 |
| | | ΔΕ/Ε | 5×10 ⁻⁴ | <5×10 ⁻⁴ @60% core (1.2nC) |
| 0.85 | 2 | emittance (mm.mrad) | 5-6 | ~6 |
| | | | | <3.5×10 ⁻⁴ @90% core (0.9nC) |
| | | ΔΕ/Ε | 5×10 ⁻⁴ | |
| | 1 | emittance (mm.mrad) | ~ 3 | ~2.3 |
| | | ΔΕ/Ε | 5×10-4 | 4×10 ⁻⁴ @80% core (1.6nC) |
| 1.4 | 2 | emittance (mm.mrad) | ~5-6 | ~4 |



A. Fedotov et al.



RF gun approach using 703MHz gun of ERL under construction at BNL (D. Kayran et al.)



20 electron bunches:

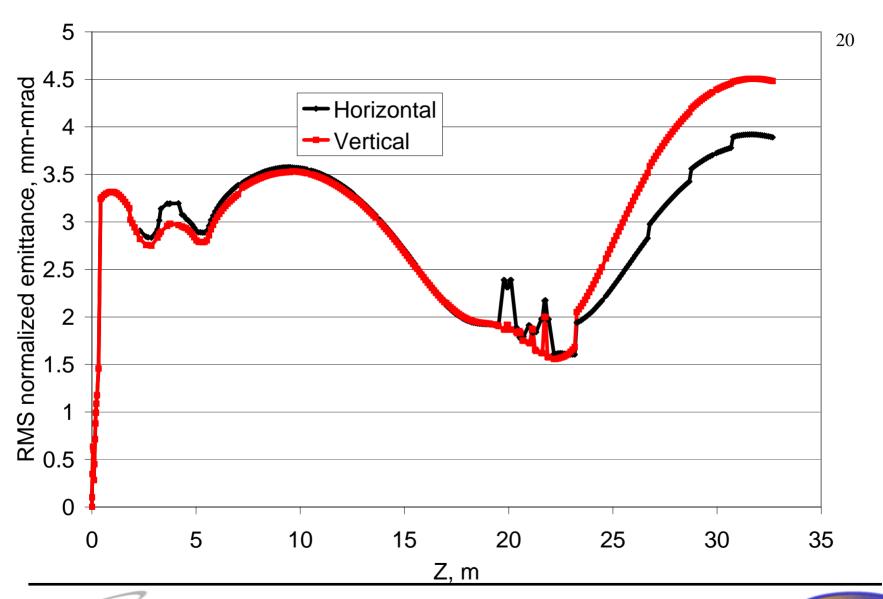
charge per bunch 50 pC, electron bunches are 42.61 cm apart

Ion bunch full bucket: length ~ 10 m





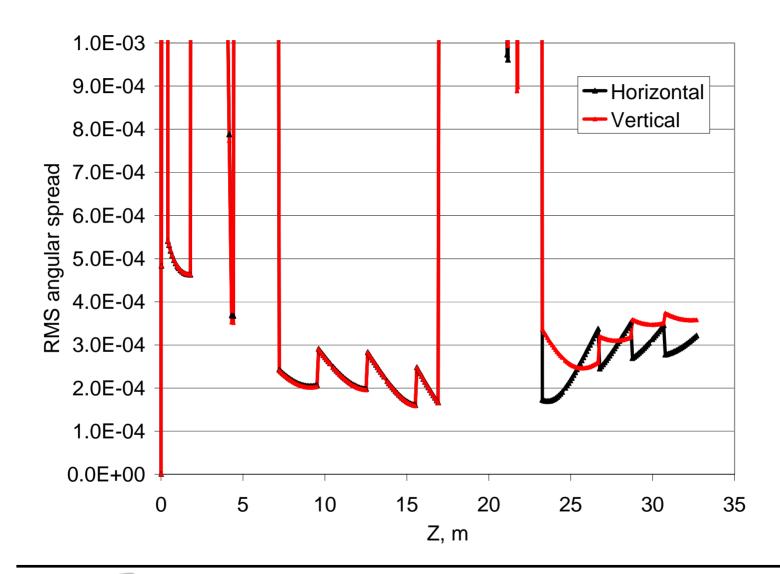
Normalized emittances evolution (for 703 MHz gun approach)





A. Fedotov et al.

RMS angular spread evolution (for 703 MHz gun approach)





A. Fedotov et al.

Possible luminosity improvements at low-energies

1. E-cooler:

• Expect factor of 3-6 improvement in L if space-charge limit is reached. Larger factor if not space-charge limited (see Ref. [6] for details)

2. Top-off mode:

- Expect factor of 2-3 improvement in L
- Replace 1-4 RHIC bunches every AGS cycle, beam stays in RHIC only 3-7 minutes.
- Needs modification of RHIC injection and extraction system
- 3. IBS below transition and use of 56 MHz cavity:
- Expect about factor of 2 in L
- 56 SRF cavity will be commissioned in FY11.





ACKNOWLEDGMENTS

We would like to thank T. Roser, W. Fischer and other members of Collider-Accelerator department at BNL for many useful discussions and suggestions.

We also thank A. Sidorin (JINR, Dubna) and S. Nagaitsev, A. Shemyakin, L. Prost (FNAL) for providing useful information.

We are grateful to Al Pendzick for providing cost estimate of possible cooling scenarios {see Ref. [6]}.





- [1] Proc. of Workshop "Can we discover QCD critical point at RHIC?" (BNL, March 2006) RIKEN BNL Research Center Report No. BNL-75692-2006; http://www..bnl.gov/riken/QCDRhic.
- [2] A. Cho, Science, V. 312, April 12, 2006, p 190.
- [3] G. Stephans, "critRHIC: the RHIC low energy program", J. Phys. G: Nucl. Part. Phys. 32 (2006).
- [4] M. Stephanov, K. Rajagopal, and E. Shuryak, Phys. Rev. Letters 81, p. 4816 (1998).
- [5] T. Satogata et al., "RHIC challenges for low energy operation", Proceed. of PAC07 (Albuquerque, NM, 2007), p. 1877; T. Satogata et al., 2008 RHIC retreat, March 31, 2008.
- [6] A. Fedotov et al., "Feasibility of electron cooling for low-energy RHIC operation", Collider-Accelerator Tech Note C-A/AP/ (April 2008).



